

Network Methods for Big Data: Applications to Astroinformatics

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- Importance of **Multiscale** approach in analysis of Astronomical Images

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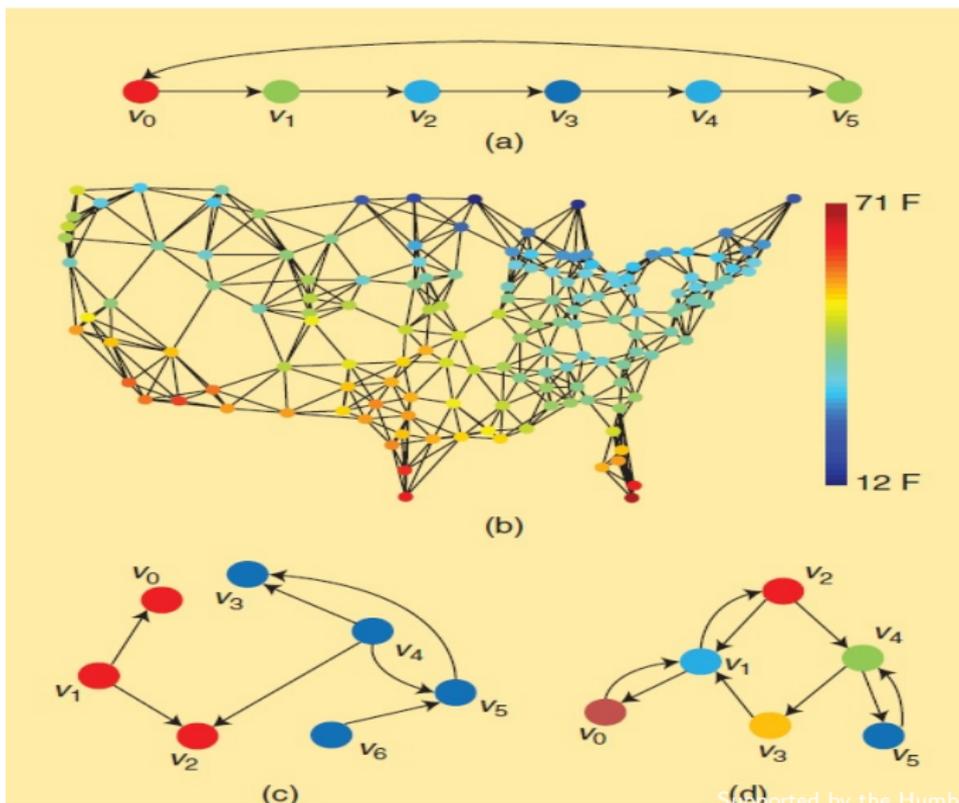
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- Challenges: Large-scale **filtering** and **frequency analysis**
- The **new notions** which generalize those of the classical DSP: **Graph signals**, Graph **filters**, Graph **Fourier transform**, Graph **frequency**, Spectrum **ordering**

3 examples of Signals on graphs

Top: Samples of the signal $\cos(2\pi n/6)$: edges show causality of time;

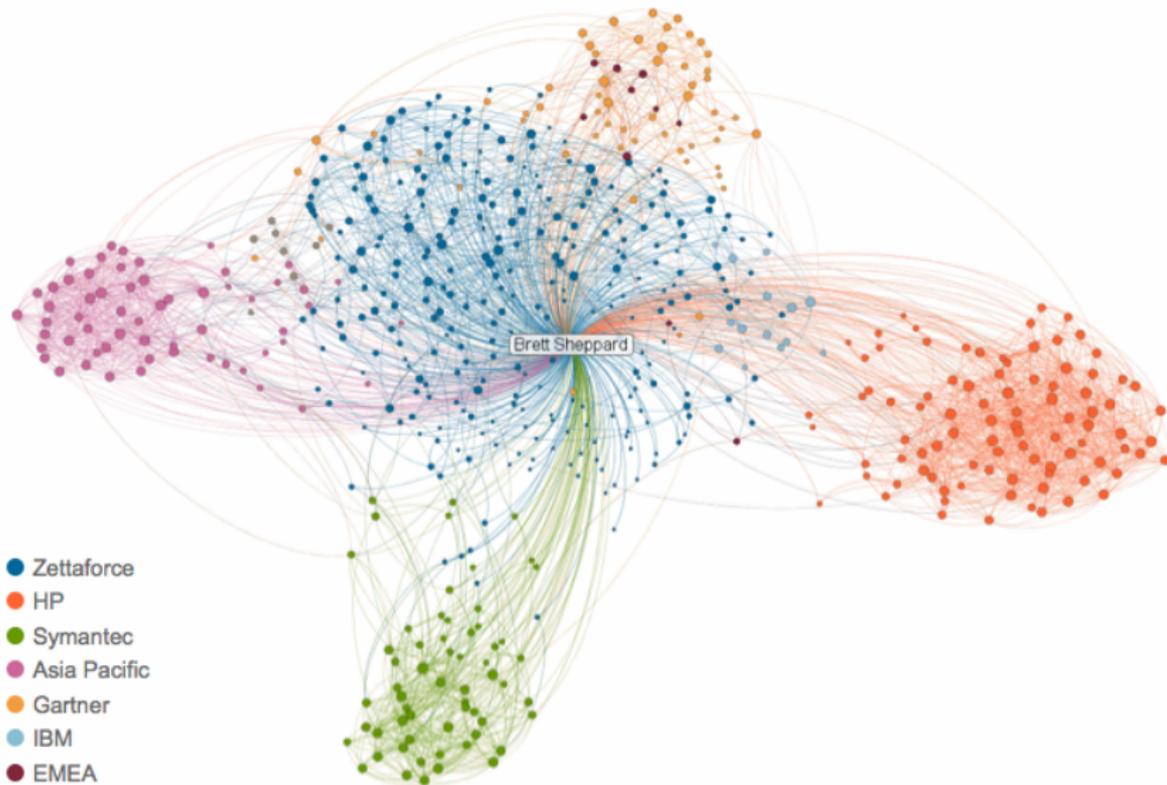


Examples above:

- **Sensor network** – edges show closeness of locations of the sensors;
- **WWW**: nodes = websites with website features, edges = hyperlinks
- **Social Network**: nodes = individuals, edges = connection and strength
- **Astronomical Images** and their Network approximations
- **Network Approximations** have **nodes** which are galaxies, stars, clusters; nodes are intersection of lines along **edges**; closer Nodes represent **similar features** within the data
- **An Exemplary OBJECTIVE of OUR WORK: Find a method for high speed automatic classification of galaxies**
- **previously done by thousands of volunteers**
- Application to: **images of distant clusters of galaxies that contain several different types of galaxies**
- General approach of **Machine Learning**: This is the art of "teaching" the algorithms the experience of people accumulated for a long time.
- Applications of the **same methods** e.g. in **medicine**, for helping to spot tumours, or in **security**, to find suspicious items in airport scans.

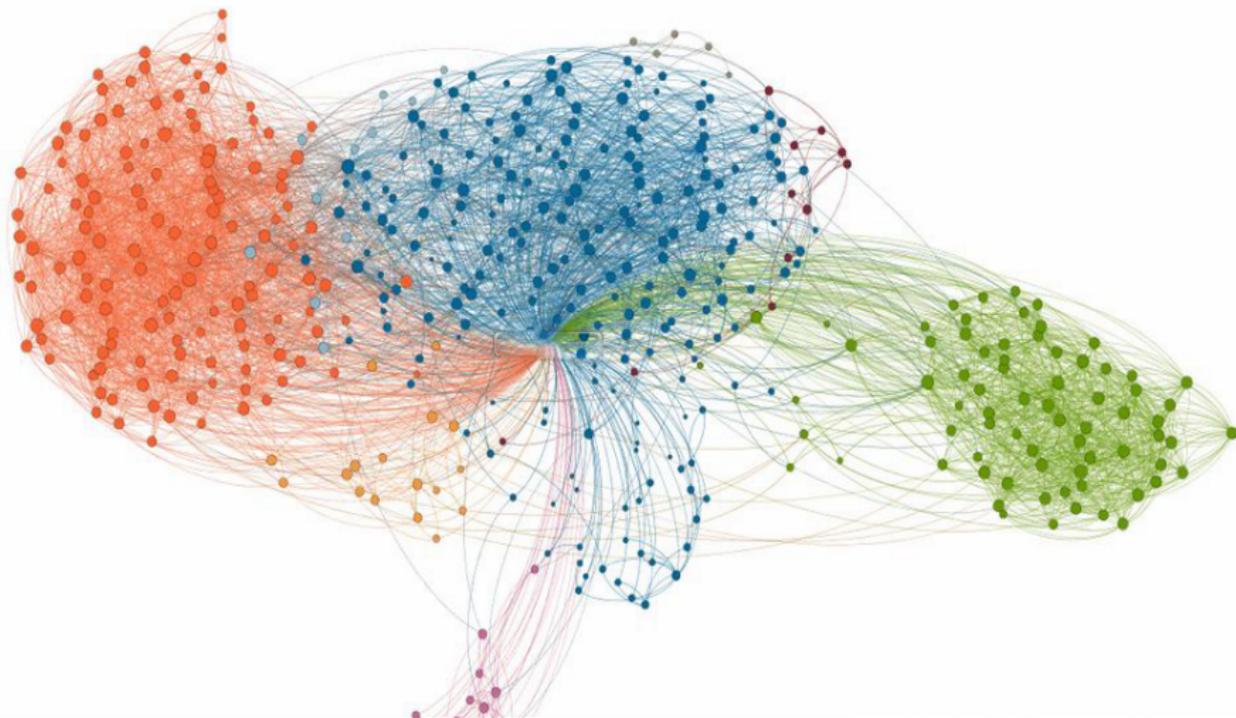
Example of Network - LinkedIn

LinkedIn Maps Brett Sheppard's Professional Network
as of January 26, 2011

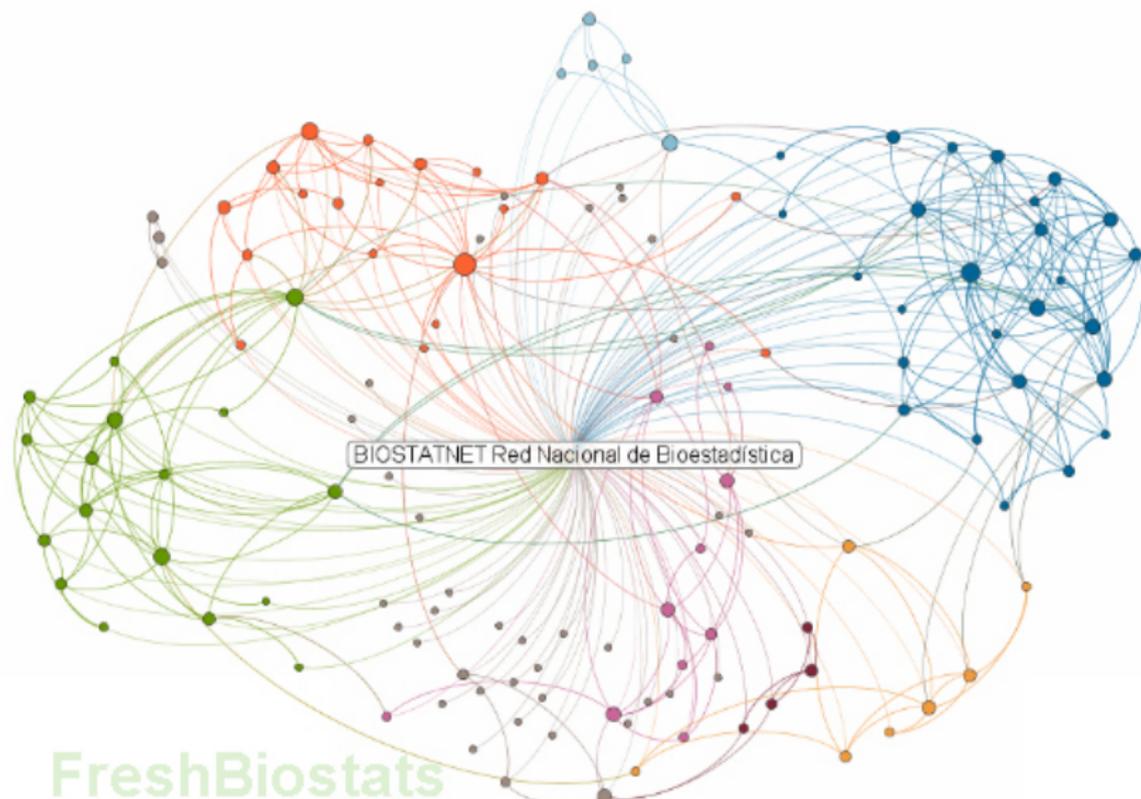


Example - LinkedIn

LinkedIn Maps of Professional Network



Biostatistics - example graph



FreshBiostats

Influencer Network-graph



SNA as a Service



General principles of SP on graphs

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- The "neighbouring" nodes for the node v_j is the set

N_j

Multiscale structure Analysis of Network Signals

- 1 Graph shift - generalizes the usual DSP **time delay**; taking average over the Neighbours

$$\bar{s}_n = \sum_{m \in N_n} A_{n,m} s_m$$

- 2 Graph Fourier transform needs the **Graph Laplace operator** (we assume undirected graph): for a signal f on the graph we define

$$Lf_m := \sum_{n \in N_m} a_{m,n} (f_m - f_n)$$

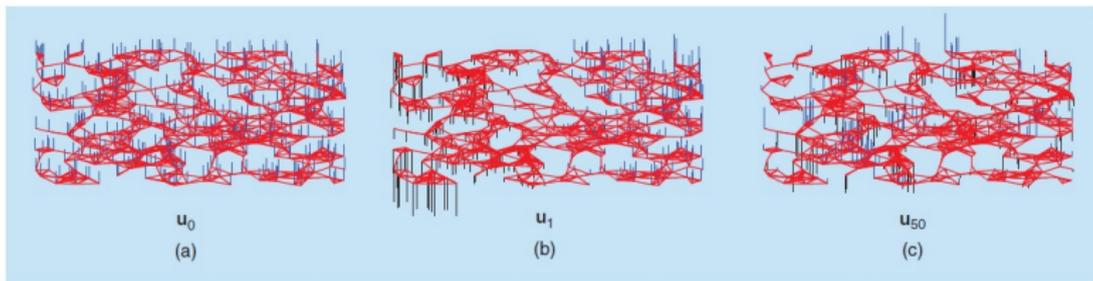
- 3 L is a real symmetric matrix: we denote the N (orthonormal) **eigenvectors and eigenvalues** of L by

$$\chi^{(n)} = (\chi_1^{(n)}, \dots, \chi_N^{(n)}) \in \mathbb{R}^N,$$

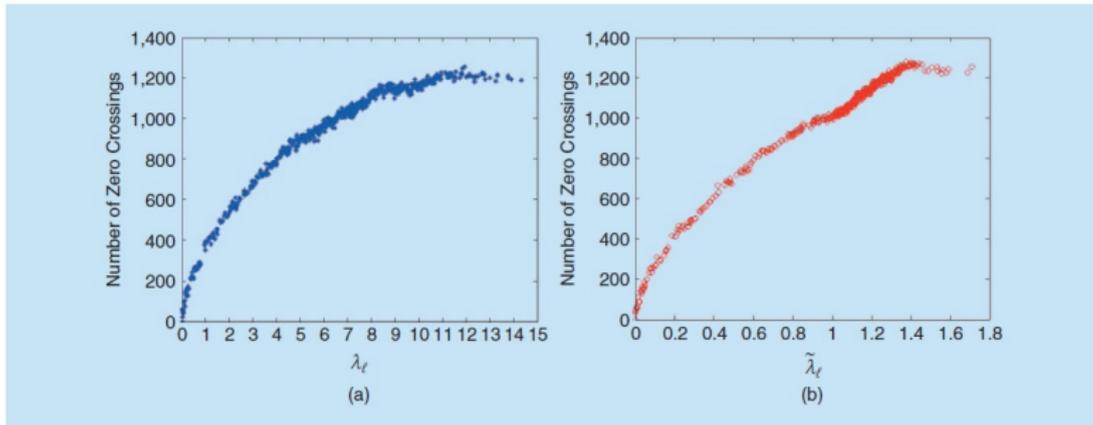
$$\lambda_0 \leq \lambda_1 \leq \dots \leq \lambda_{N-1}.$$

They are the analogue to e^{ikt} , since $-\frac{d^2}{dt^2} e^{ikt} = k^2 e^{ikt}$

Example of eigenvectors



[FIG2] (a)–(c) Three graph Laplacian eigenvectors of a random sensor network graph. The signals' component values are represented by the blue (positive) and black (negative) bars coming out of the vertices. Note that u_{50} contains many more zero crossings than the constant eigenvector u_0 and the smooth Fiedler vector u_1 .



[FIG3] The number of zero crossings, $|\mathcal{Z}_G(u_\ell)|$ in (a) and $|\mathcal{Z}_G(\tilde{u}_\ell)|$ in (b), of the unnormalized and normalized graph Laplacian eigenvectors for the random sensor network graph of Figure 2, respectively (the latter of which is defined in the section "Other Graph

Fourier and Wavelet Analysis on Graphs

- Hence, we have Fourier Analysis

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- The role of t - the **scale** of the localized at **node** a wavelet $\psi_{t,a}$ where $(a \in G)$.

Examples of localizations – Swiss roll

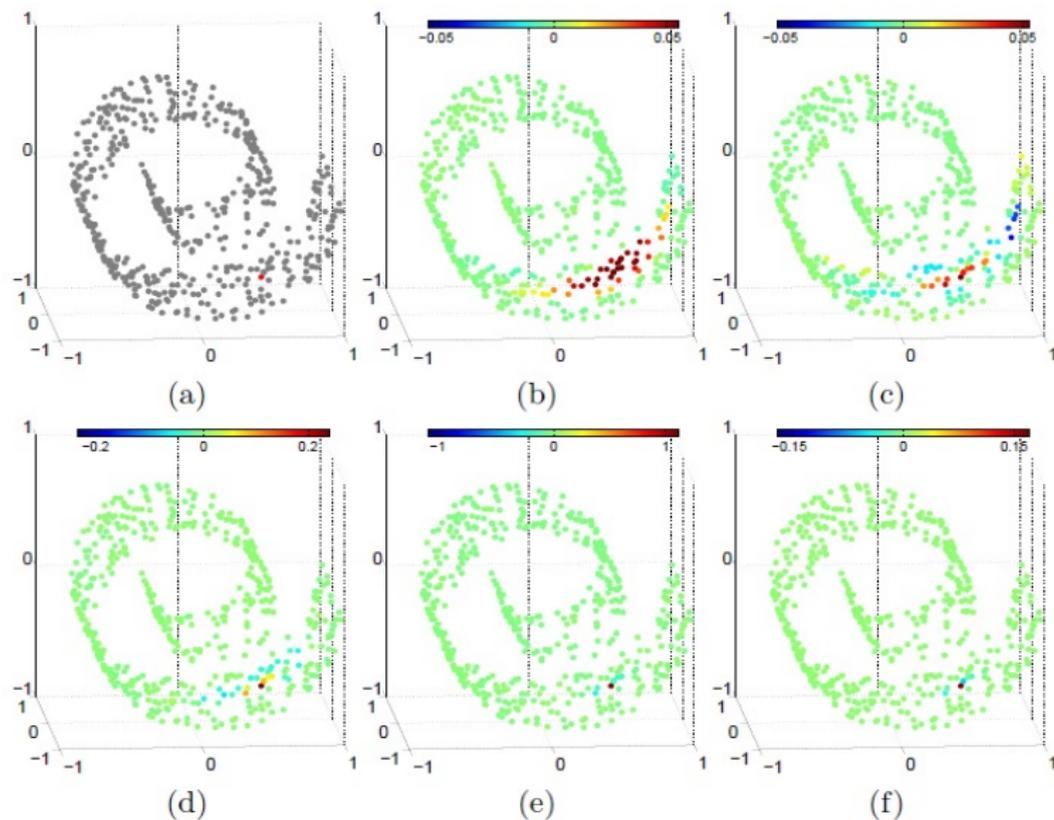


Figure 2: **Geometric localization on a Swiss Roll dataset**, courtesy of The Numerical Simulation of

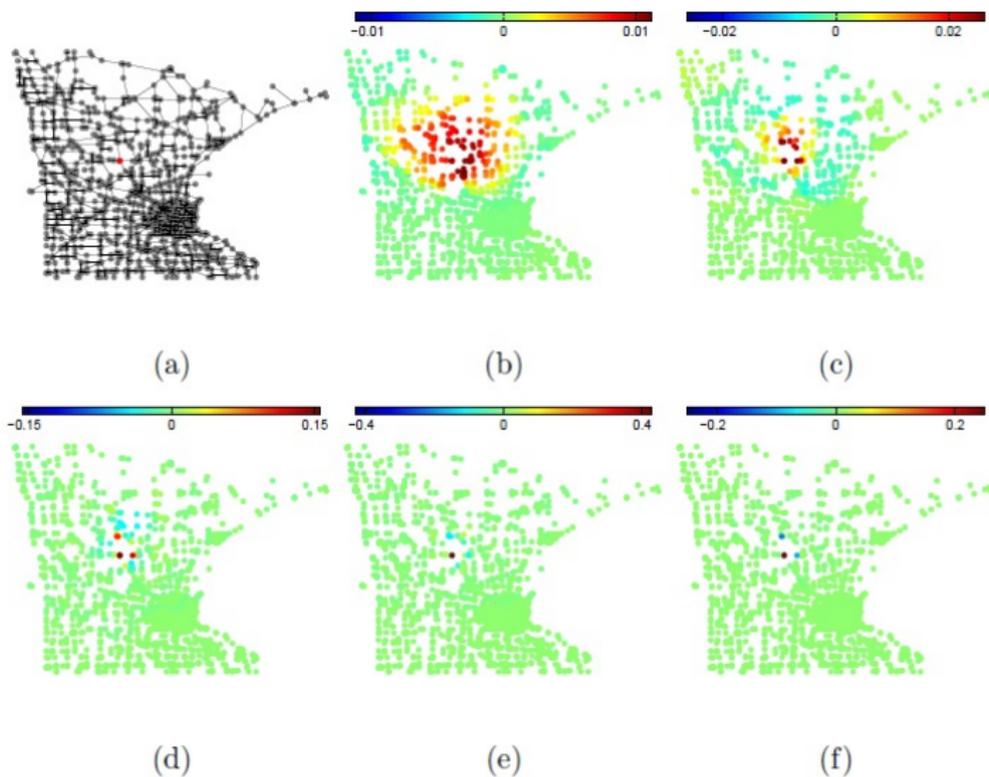


Figure 4: Spectral graph wavelets on Minnesota road graph, with $K = 100$, $J = 4$ scales. (a) vertex at which wavelets are centered (b) scaling function (c)-(f) wavelets, scales 1-4.

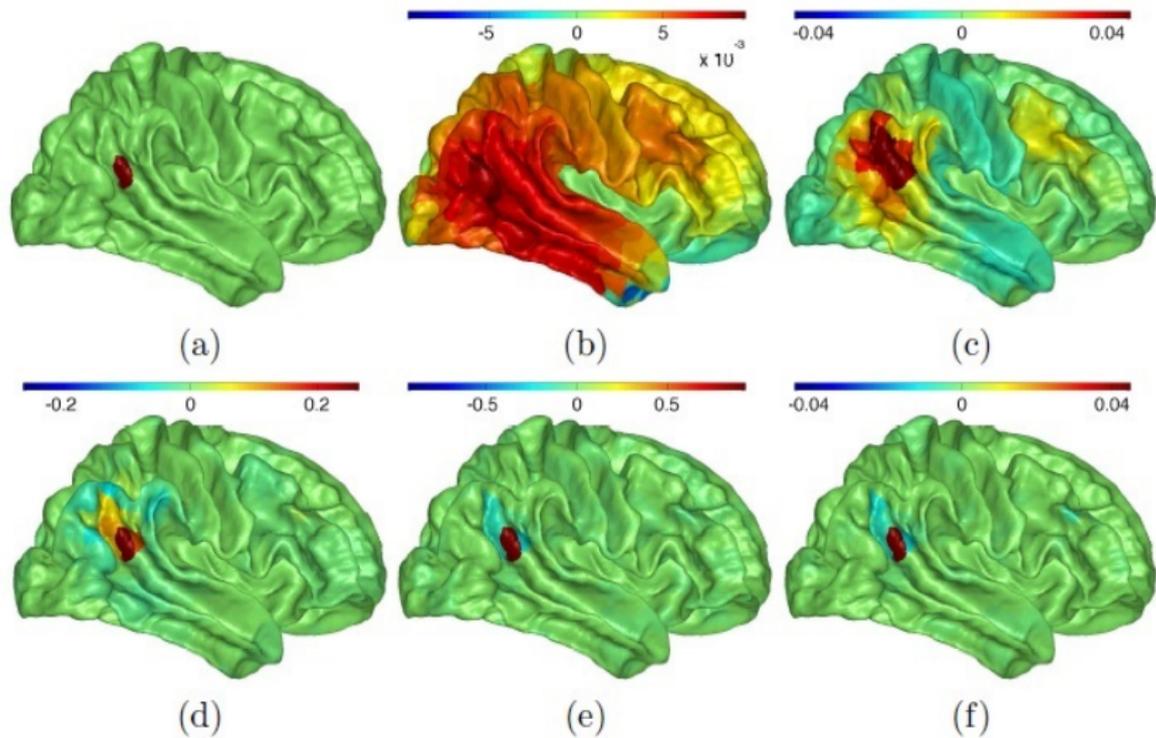


Figure 5: Spectral graph wavelets on cerebral cortex, with $K = 50$, $J = 4$ scales. (a) ROI at which wavelets are centered (b) scaling function (c)-(f) wavelets, scales 1-4.

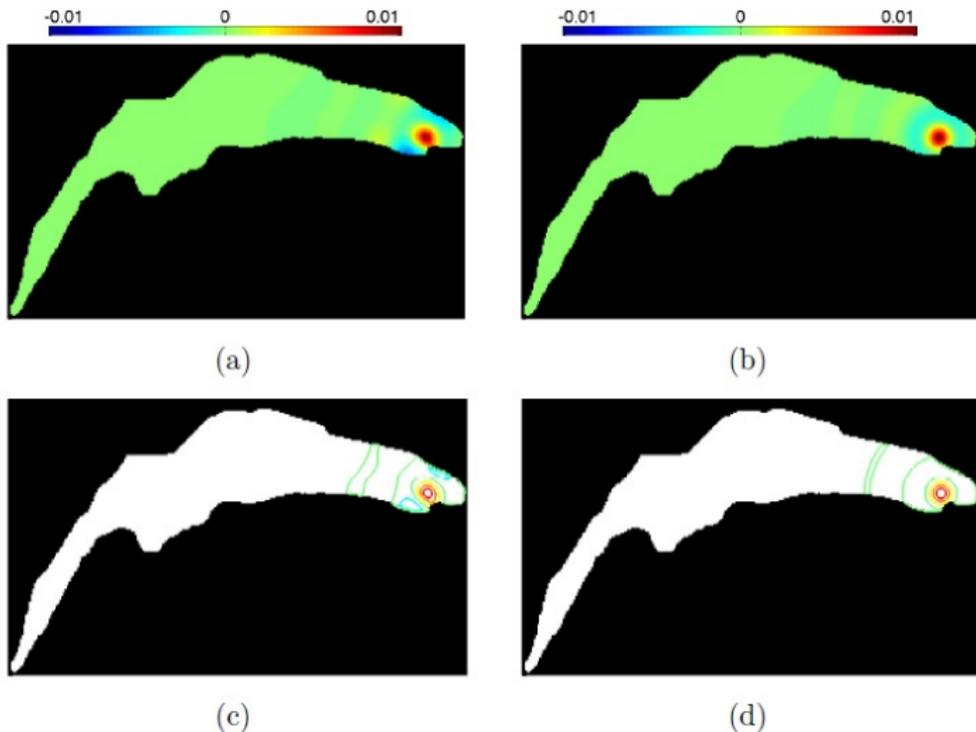
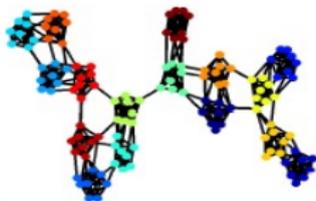


Figure 6: Spectral graph wavelets on lake Geneva domain, (spatial map (a), contour plot (c)); compared with truncated wavelets from graph corresponding to complete mesh (spatial map (b), contour plot (d)). Note that the graph wavelets adapt to the geometry of the domain.

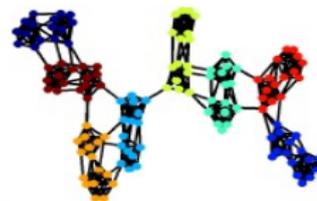
Main Objective: Clusterization (Community Detection on graphs)

Multiscale community structure in a graph

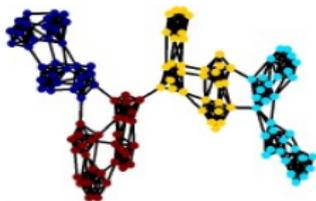
finest scale (16 com.):



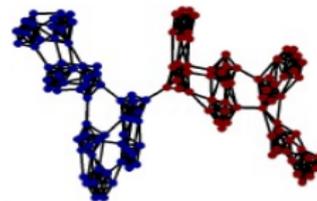
coarser scale (8 com.):



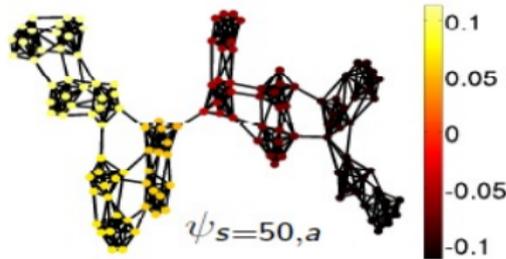
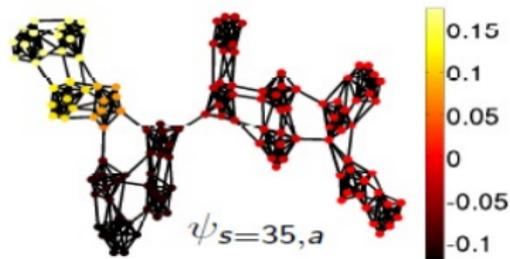
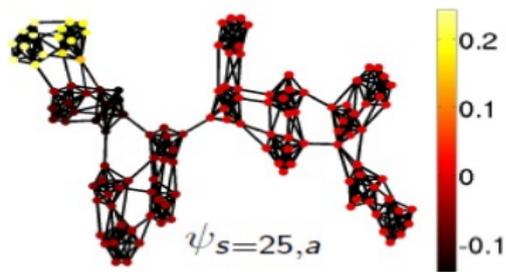
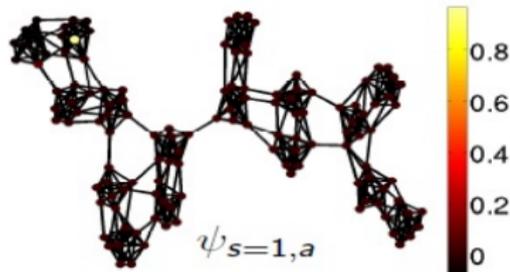
even coarser scale (4 com.):



coarsest scale (2 com.):

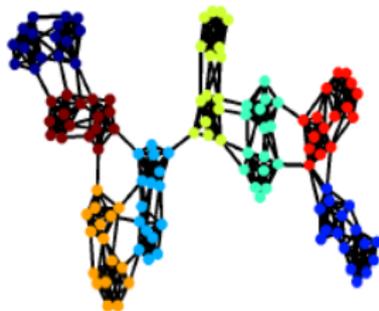


Examples of wavelets



Comparison with other methods for Community Detection

Classical community detection algorithms do not have this “scale-vision” of a graph. Modularity optimisation finds:



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- **Astronomical image segmentation by Networks and wavelets**

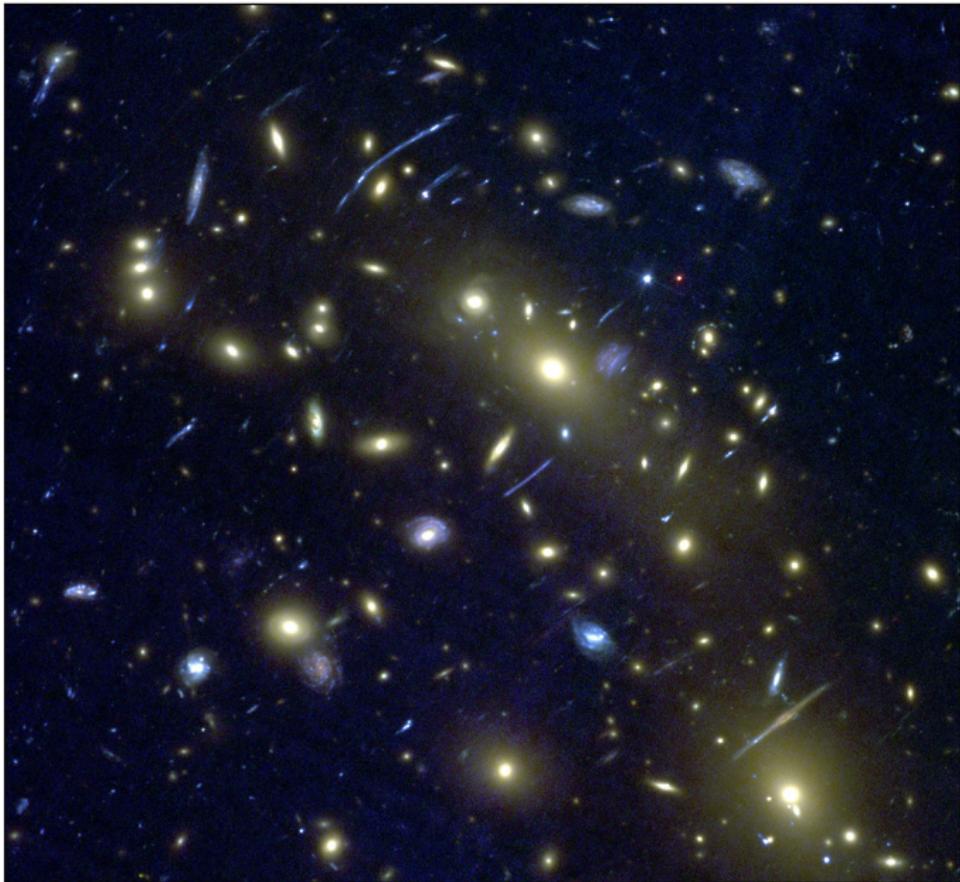
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- Astronomical image segmentation by Networks and wavelets
- Machine learning method to analyse galaxy images

Pictures to:

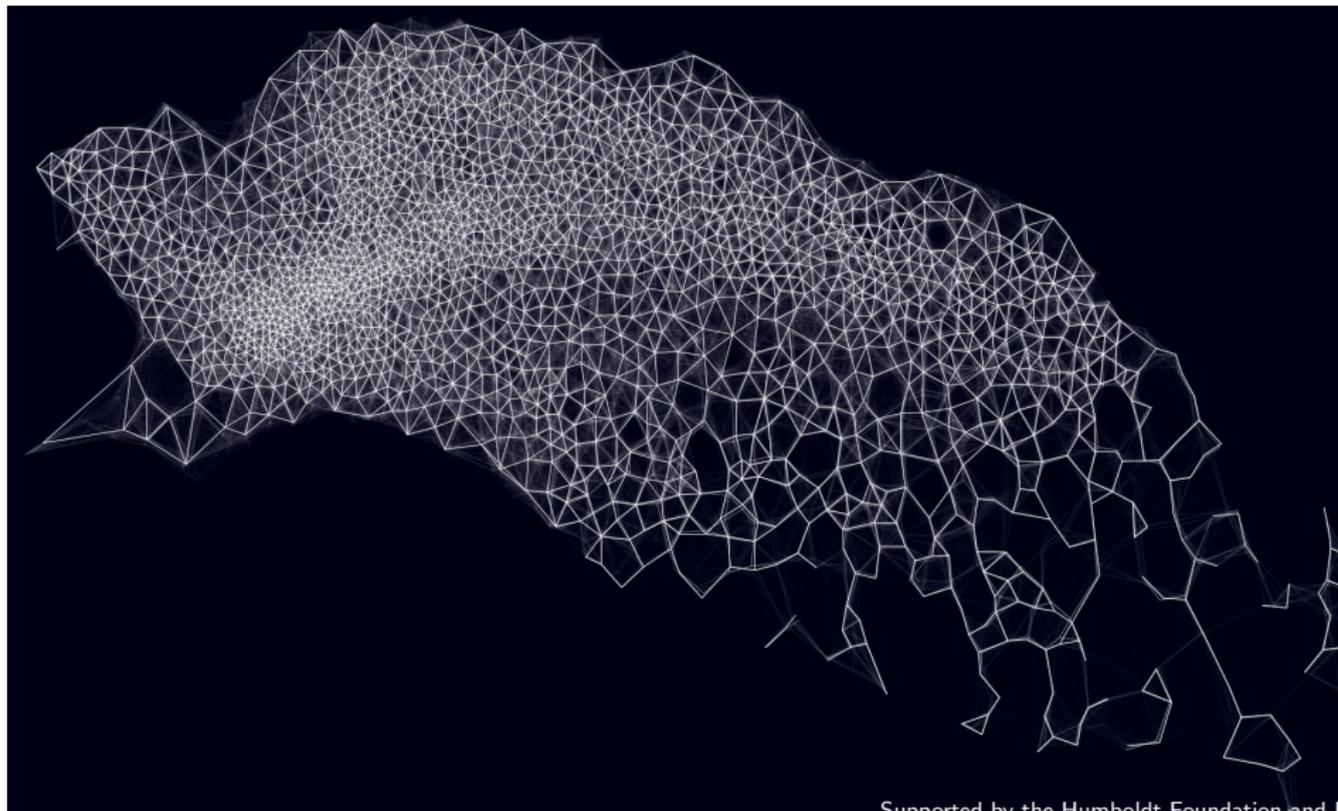
Machine learning: to "teach" a machine to analyse galaxy images

- 1 Picture 1: Hubble Space Telescope image of the **cluster** of galaxies MACS0416.1-2403, one of the Hubble 'Frontier Fields'. Bright yellow 'elliptical' galaxies can be seen, surrounded by numerous blue **spiral** and amorphous (star-forming) galaxies
- 2 Picture 2: Visualisation of the (**neural**) **network** representing the 'brain' of the machine learning algorithm. The intersections of lines are called nodes, and these represent a map of the input data. Nodes that are closer to each other represent similarity
- 3 Picture 3: A **zoom-in of part** of the network described above.
Credit: J. Geach / A. Hocking
- 4 Picture 4: Image showing the MACS0416.1-2403 **cluster**, highlighting parts of the image that the algorithm has identified as '**star-forming**' galaxies. Credit: NASA / ESA / J. Geach / A. Hocking
- 5 Picture 5: Image showing the MACS0416.1-2403 **cluster**, highlighting parts of the image that the algorithm has identified as '**elliptical**' galaxies. Credit: NASA / ESA / J. Geach / A. Hocking

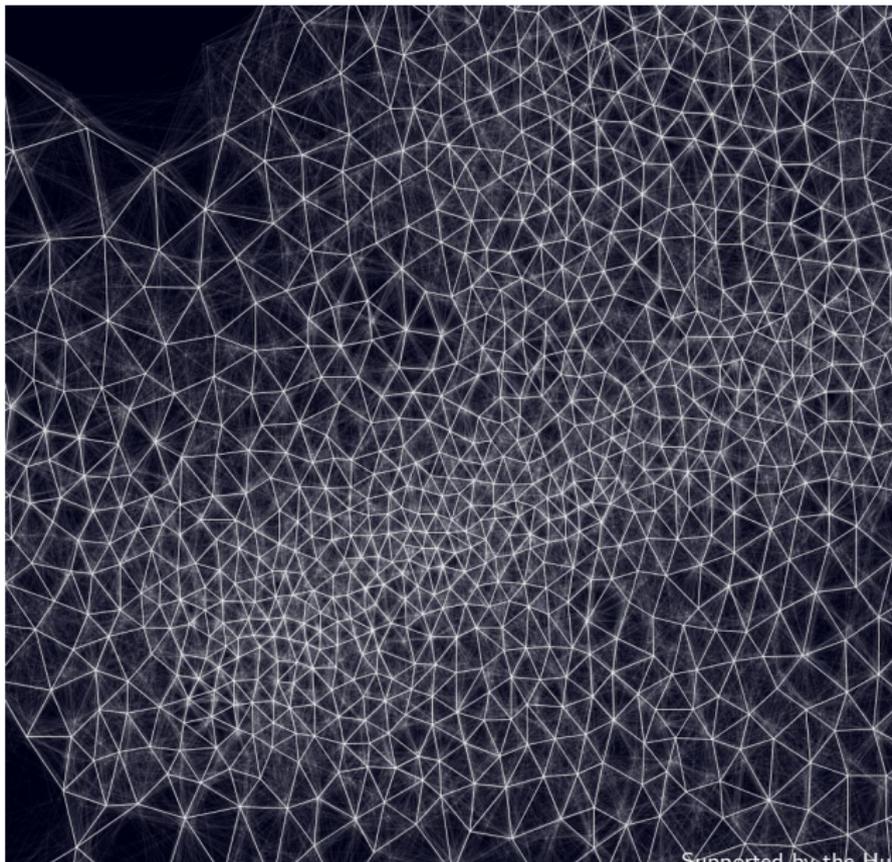




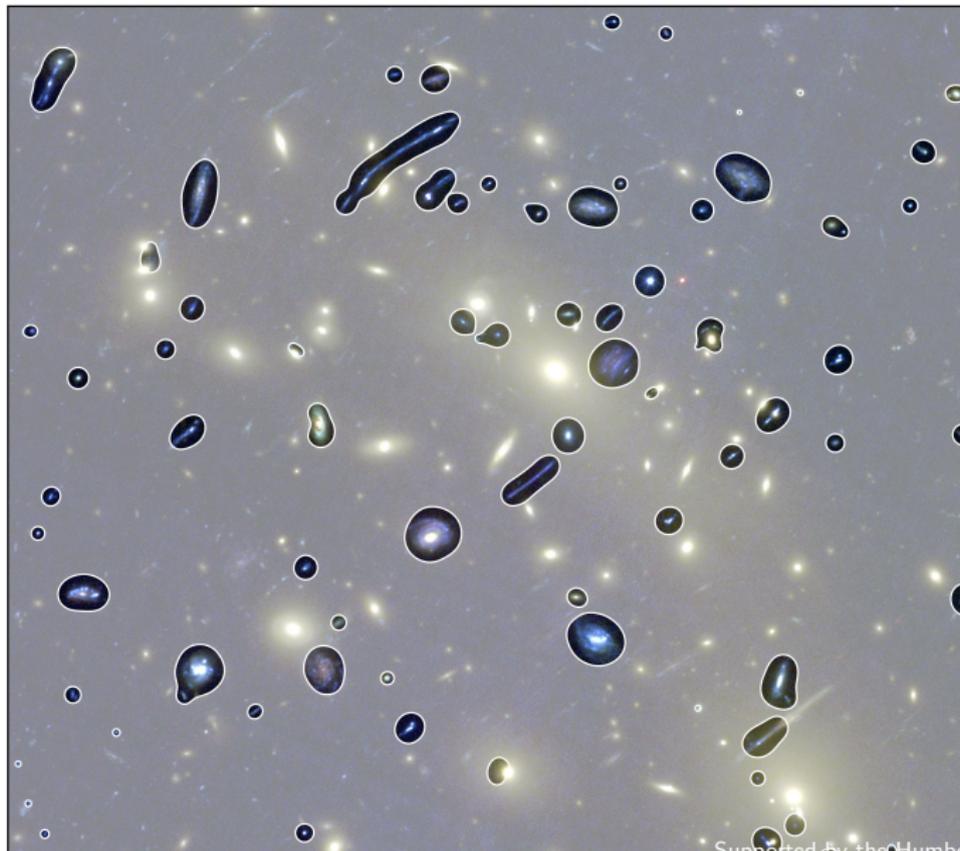
Continued - Network approximation



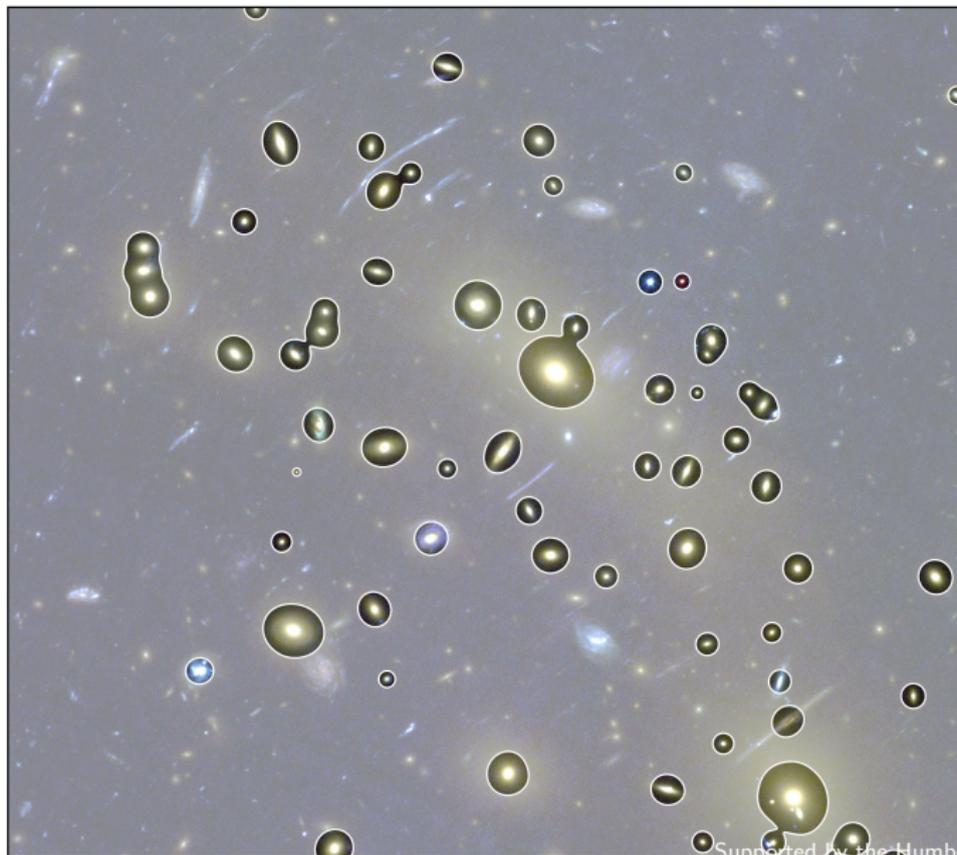
Cont'd - zoomed part



Cont'd - "star forming" galaxies



Cont'd – "elliptical" galaxies



Deep Learning and Wavelets approach

References and Credits

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