

**A STUDY OF SUPERSONIC TURBULENCE  
IN STAGNATING PLASMAS**

**E. Stambulchik<sup>1</sup>, E. Kroupp<sup>1</sup>, A. Starobinets<sup>1</sup>, D. Osin<sup>1</sup>, V. I. Fisher<sup>1</sup>,  
D. Alumot<sup>1</sup>, Y. Maron<sup>1</sup>, S. Davidovits<sup>2</sup>, N. J. Fisch<sup>2</sup> and A. Fruchtman<sup>3</sup>**

<sup>1</sup>*Weizmann Institute of Science, Rehovot 7610001, Israel*

<sup>2</sup>*Princeton University, Princeton, New Jersey 08540, USA*

<sup>3</sup>*H.I.T.—Holon Institute of Technology, Holon 5810201, Israel*

*E-mail: Evgeny.Stambulchik@weizmann.ac.il, s davidov@princeton.edu*

Evolution of the ion kinetic energy in a stagnating z-pinch plasma was determined from Doppler-dominated lineshapes augmented by measurements of plasma properties and assuming first a uniform-plasma model. Notably, the energy was found to be dominantly stored in hydrodynamic flow. The Reynolds and Mach numbers are such that this motion could be supersonically turbulent, implying a non-uniform distribution of the plasma density. The data was re-analyzed under this assumption, resulting in a substantially decreased inferred mean density, while improving agreement of the model with observations. Beyond aiding our understanding of z-pinches, it is hoped that this study has highlighted fertile ground for relation to problems of astrophysical interest, such as the star formation efficiency or molecular cloud dynamics.